Overview of Fundamental Principles in Statistical Mechanics

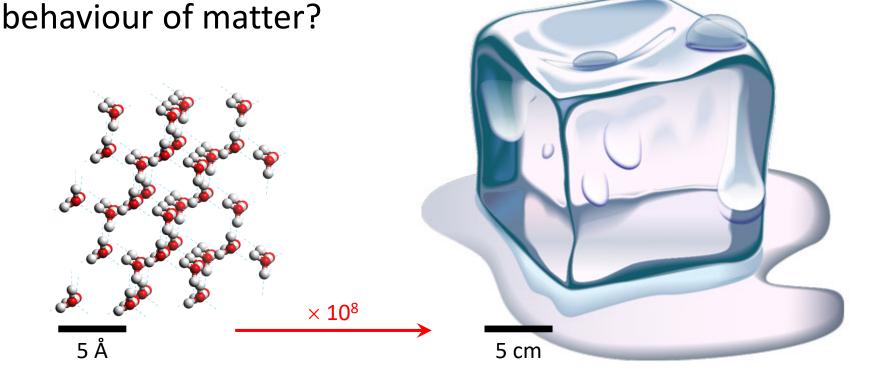
Thomas Young Centre Materials Modelling Course

Dr Michail Stamatakis

Bridging with the Microscopic World

• In the course, so far we have studied "small" entities: electrons, atoms, molecules...

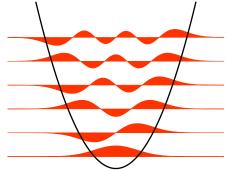
How do the laws applicable thereto shape the macroscopically observable

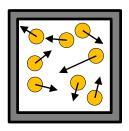


Outline

- Thermodynamics *Memorandum*
 - Refreshing our memory on basic concepts (which we will use...)
- The Statistical Mechanics Framework
 - Microstates
 - Hamiltonian
 - Averaging
- Statistical-Mechanical Ensembles
 - Microcanonical
 - Canonical
 - Grand Canonical



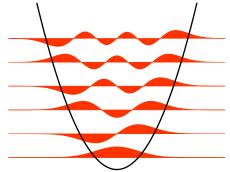


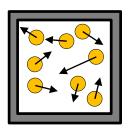


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Thermodynamic Description of Systems

 Consider a thermodynamic system, e.g. water at the triple point in a box...

"Macroscopic" description in terms of:

Intensive variables:

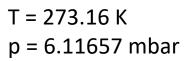
Temperature (T)

Pressure (p)

Chemical potential (µ)

Extensive variables:

Number of moles/molecules (N) Volume (V) Entropy (S)



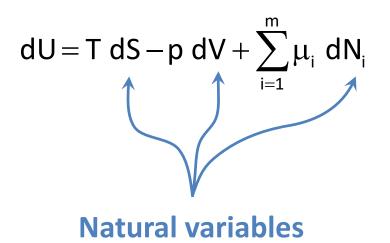
Thermodynamic potentials:

Internal/Thermal Energy (U) Enthalpy (H)

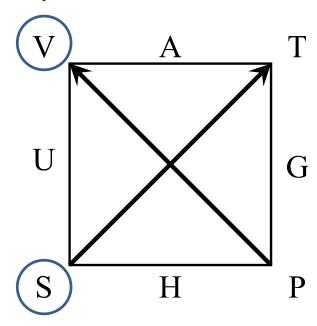
Helmholtz free energy (A) Gibbs free energy (G)

Interrelations between Thermodynamic Potentials

Interrelations between these thermodynamic variables:



$$\left. \frac{\partial U}{\partial S} \right|_{V,\{N_i\}} = T$$
 $\left. \frac{\partial U}{\partial V} \right|_{S,\{N_i\}} = -p$



Similar relations hold for the other variables...

Interrelations between Thermodynamic Potentials

Interrelations between these thermodynamic variables:

$$dU = T dS - p dV + \sum_{i=1}^{m} \mu_i dN_i$$

$$dH = T dS + V dp + \sum_{i=1}^{m} \mu_i dN_i$$

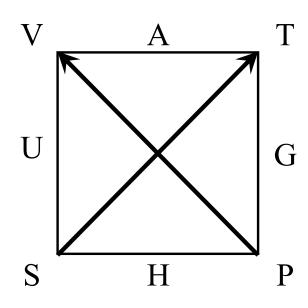
$$dG = -S dT + V dp + \sum_{i=1}^{m} \mu_i dN_i$$

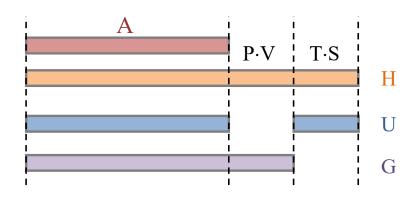
$$dA = -S dT - p dV + \sum_{i=1}^{m} \mu_i dN_i$$

$$H = U + P V$$

$$G = H - T S$$

$$A = U - TS = G - PV = H - PV - TS$$





Euler's Theorem for Extensive Variables

- Euler's theorem for first-order homogeneous functions:
 - If function f satisfies: $f(\lambda x_1, \lambda x_2, ..., \lambda x_n) = \lambda f(x_1, x_2, ..., x_n)$

- Then:
$$f(x_1, x_2, ..., x_n) = \sum_{i} \frac{\partial f}{\partial x_i} \Big|_{x_{i+2i}} x_i$$

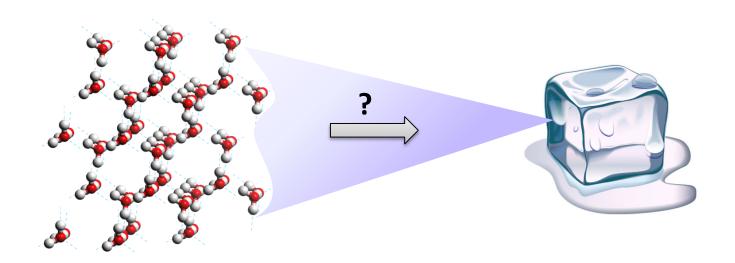
 This can be used for the thermodynamic potentials as they are indeed first-order homogeneous functions, e.g.:

$$U = \frac{\partial U}{\partial V}\bigg|_{S,\{N_i\}} V + \frac{\partial U}{\partial S}\bigg|_{V,\{N_i\}} S + \sum_i \frac{\partial U}{\partial N_i}\bigg|_{S,V} N_i$$

$$U = -pV + TS + \sum_{i} \mu_{i} N_{i}$$

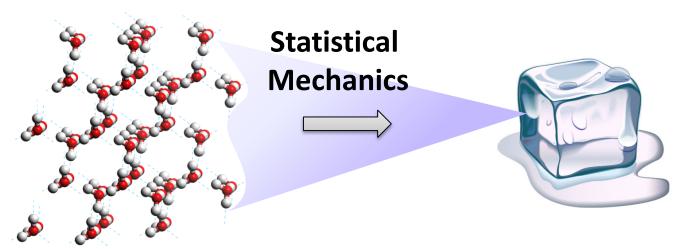
Bridging with the Microscopic World

- Within the framework of Chemical Thermodynamics we can study the interrelation of heat and work with chemical reactions or with physical changes of state...
- But how are these macroscopic processes, properties or quantities related to the "molecular world"?



The Challenge

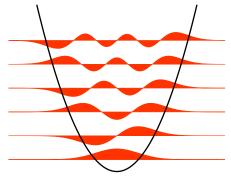
- Particles (atoms, molecules) obey known laws
 (Schrodinger equation, Newton's laws of motion) with specified inter-particle interactions...
- What are the observable properties of a system consisting of a (very) large number of such particles?
- Solving the microscopic equations for $\mathcal{O}(10^{23})$ is impossible, but we can invoke statistical arguments...

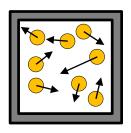


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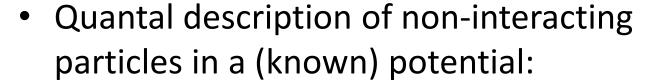


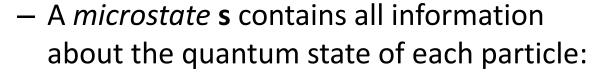


Microstates

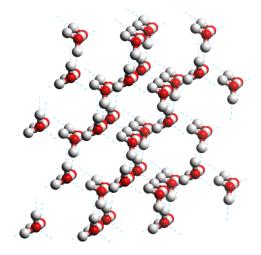
- Key to these statistical arguments: how frequently a microstate is visited... but what is a microstate?
- Classical description of an ice-cube:
 - A microstate s contains all information about the position and momentum of each atom:

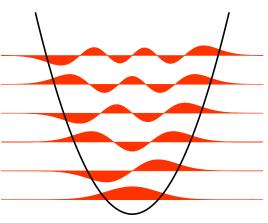
$$s = (r_1, r_2, ..., r_N; p_1, p_2, ..., p_N)$$





$$\mathbf{s} \equiv (\mathbf{n}_1, \mathbf{n}_2, ..., \mathbf{n}_N)$$

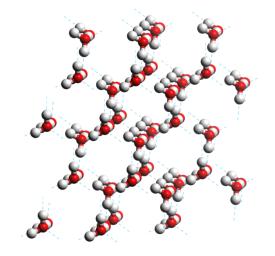




Hamiltonian

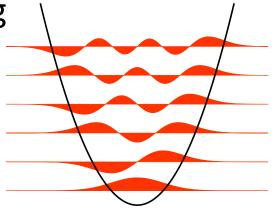
- The Hamiltonian of a system gives the energy of a microstate
- Classical Hamiltonian of an ice-cube:

$$\mathcal{H}(\mathbf{s}) = \sum_{i=1}^{N} \frac{\left|\mathbf{p}_{i}\right|^{2}}{2 m_{i}} + \mathcal{V}(\mathbf{r}_{1}, \mathbf{r}_{2}, ..., \mathbf{r}_{N})$$

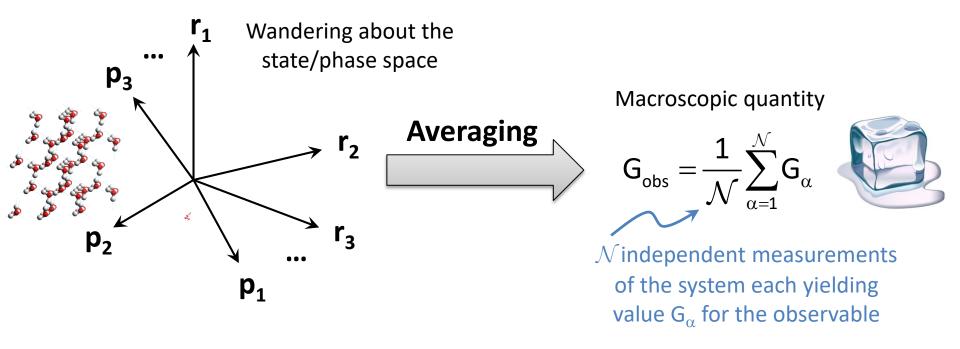


 Quantal Hamiltonian of non-interacting particles in a (known) potential:

$$\mathcal{H} = -\frac{\hbar^{2}}{2} \sum_{i=1}^{N} \frac{1}{m_{i}} \nabla^{2} + \mathcal{V}(\mathbf{r_{1}}, \mathbf{r_{2}}, ..., \mathbf{r_{N}})$$



From Micro to Macro via Averaging

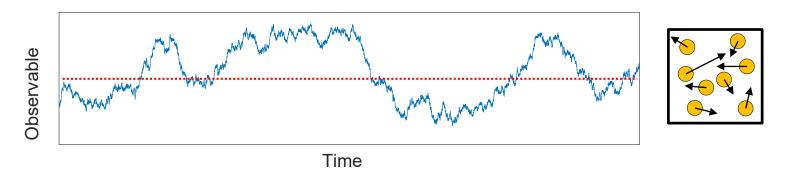


- Fundamental postulates in Statistical Mechanics:
 - Principle of equal weights: microstates with equal energy are visited with the same frequency
 - Ergodic hypothesis: averages over time are equal with ensemble averages...

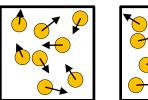
Averaging Operations and Ensembles

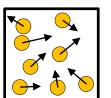
Let us consider the time average of an observable...
 and try to partition the sum as follows:

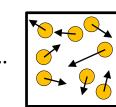
$$G_{obs} = \sum_{\alpha=1}^{N} \sum_{\alpha=1}^{N} G_{\alpha} = \sum_{\alpha} \frac{1}{N} \begin{cases} \text{Number of times state } \mathbf{s}_{\alpha} \text{ is observed in the } \mathcal{N} \text{ samples} \end{cases} G_{\alpha}$$
Time average...



$$\Rightarrow G_{obs} = \sum_{\chi} P_{\chi} G_{\chi} = \langle G \rangle$$
 Ensemble average...





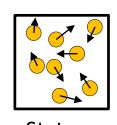


"Copies" of a system, each in a potentially different micro-state...

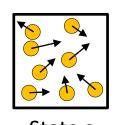
Ensembles

$$G_{obs} = \sum_{\chi} P_{\chi} G_{\chi} = \langle G \rangle$$

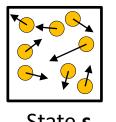




State **s**₁ Probability P₁



State **s**₂ Probability P₂



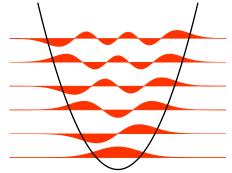
State \mathbf{s}_{χ} Probability P_{χ}

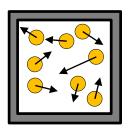
- An ensemble is the collection/assembly of all possible microstates, which the system can visit given some macroscopic constraints, for example:
 - MICROCANONICAL ENSEMBLE: all states with fixed number of molecules, volume and energy (N, V, E)
 - CANONICAL ENSEMBLE: all states with fixed number of molecules, volume and temperature (N, V, T)
 - GRAND CANONICAL ENSEMBLE: all states with fixed
 chemical potential, volume and temperature (μ, V, T)

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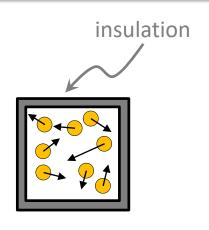




Microcanonical Ensemble (N, V, E)

- All microstates have the same energy ⇒
 they are visited with equal probability
- For discrete systems, define the degeneracy as:

$$\Omega(N,V,E) = \begin{pmatrix} Number of microstates \\ with N, V and energy E \end{pmatrix}$$



No matter exchange No heat exchange

 Analogously, for continuous systems one can define the density of states as:

$$\overline{\Omega}(N,V,E)dE = \begin{pmatrix} Number of microstates with \\ N, V and energy between E, E+dE \end{pmatrix}$$

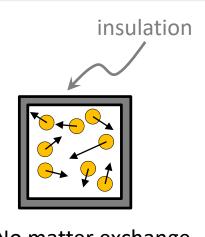
Microcanonical Ensemble

The probability of each microstate is:

$$P_{\chi} = \frac{1}{\Omega(N,V,E)}$$



- States that are not in the ensemble have probability of zero.
- For continuous systems P_{χ} is replaced with a probability density, but similar equations hold. To keep the discussion simple we will focus on the discrete case.



No matter exchange No heat exchange

How does this help us understand the macroscopic world

Definition of Entropy

 The microcanonical ensemble allows us to define entropy as:

$$S = k_B \log(\Omega(N,V,E))$$

where k_B is Boltzmann's constant:

$$k_B = 1.38064852 \times 10^{-23} \text{ J} \cdot \text{K}^{-1}$$

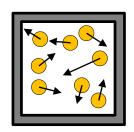
= 8.6173303 × 10⁻⁵ eV·K⁻¹

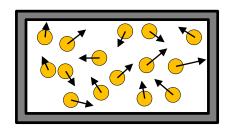


Ludwig Eduard Boltzmann

• By the above definition entropy is **extensive**: if a system is composed of two independent subsystems A and B:

$$\begin{split} &\Omega_{\text{AB}} = \Omega_{\text{A}} \cdot \Omega_{\text{B}} \implies \\ &S_{\text{AB}} = k_{\text{B}} \log \left(\Omega_{\text{A}} \cdot \Omega_{\text{B}} \right) = S_{\text{A}} + S_{\text{B}} \end{split}$$





Temperature in the Microcanonical Ensemble

 Since all microstates have the same energy in the microcanonical ensemble, the thermal energy is: U = E

• Recall that:
$$T = \frac{\partial U}{\partial S}\Big|_{V,N}$$
 (see slide 5)

• Therefore, the temperature can be calculated from:

$$\beta = (k_B \cdot T)^{-1} = \frac{\partial \log(\Omega)}{\partial E} \bigg|_{V,N}$$

- β > 0 for macroscopic systems encountered in nature (for which a classical description is adequate)
- For certain quantum systems β can become negative (see e.g. population inversion in laser physics).

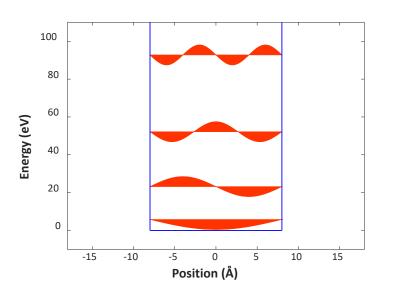
Pressure in the Microcanonical Ensemble

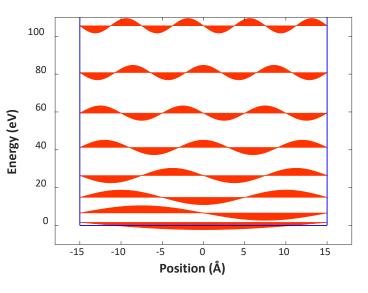
• Since U = E, the pressure is:

$$p = -\frac{\partial E}{\partial V}\Big|_{S.N}$$
 (see slide 5)

- The volume affects the energy levels!
 - e.g. for a quantum particle in a 1D box:

$$E_n = \frac{n^2 h^2}{8 m L^2}$$



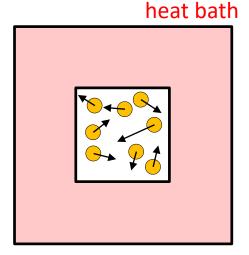


 Now that we have S, p, U and T we can calculate any other thermodynamic quantity (see slides 5, 6)

Canonical Ensemble (N, V, T)

- This ensemble comprises microstates with fixed volume and number of molecules
- Energy can fluctuate and the system is at thermal equilibrium with a heat bath kept at temperature T

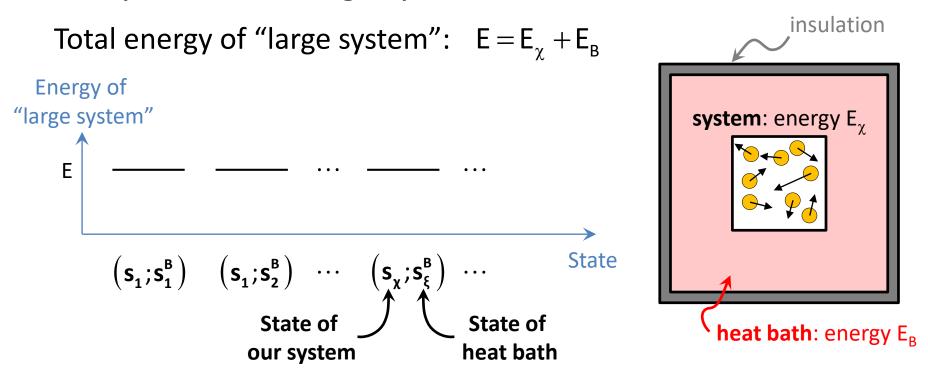




No matter exchange Heat exchange

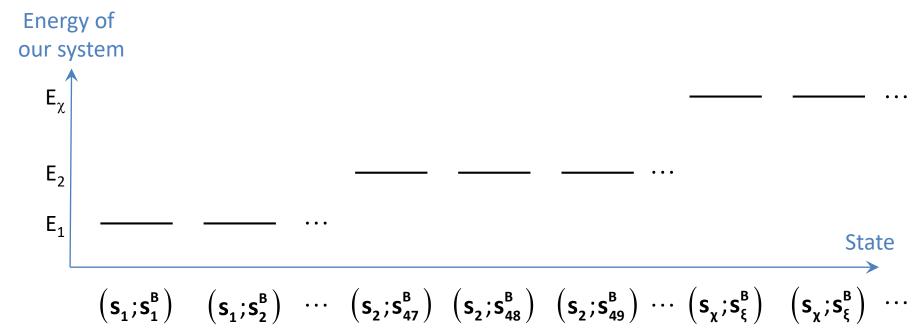
What is the probability that each microstate is visited in this ensemble

 Consider the system in the canonical ensemble as a subsystem of a "large system" in the microcanonical



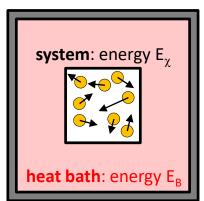
 Proper combinations of these microstates (i.e. with energies summing up to E) are indeed valid (accessible) states of the large system in the microcanonical ensemble...

Consider now a different partitioning of these states:

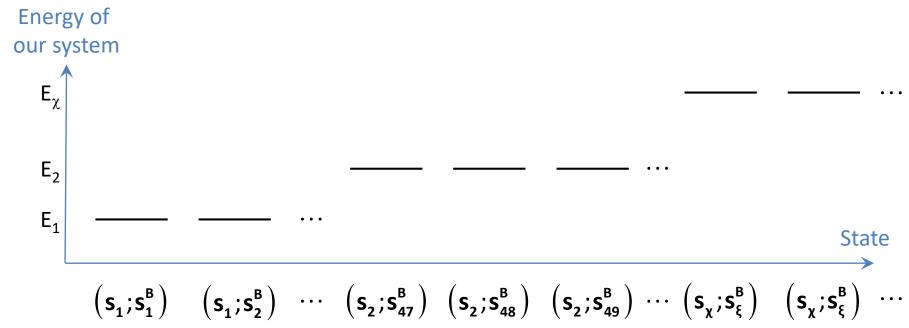


• For a moment, fix the system in state χ . Then the accessible states for the large system are those of the bath, hence:

$$\Omega(E_B) = \Omega(E - E_\chi)$$
 since $E = E_\chi + E_B$

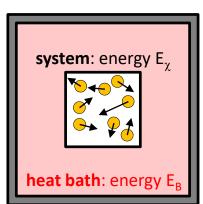


Consider now a different partitioning of these states:

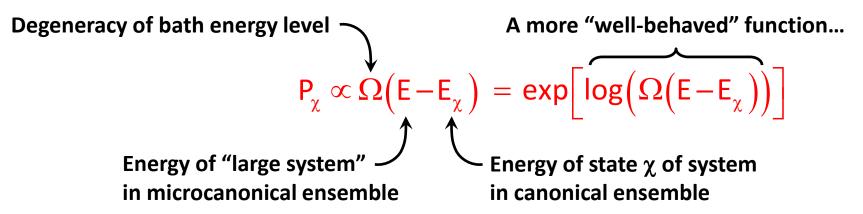


• By the *principle of equal weights* the probability of state χ of the system is:

$$P_{\chi} \propto \Omega(E_{B}) = \Omega(E-E_{\chi})$$

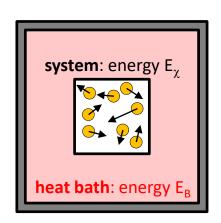


• So, the probability of state \mathbf{s}_{χ} in the canonical ensemble:



- The presence of Ω and E in this equation is inconvenient!
- Since E_{χ} is small, introduce a Taylor expansion for $\log(\Omega)$:

$$\begin{split} \log \left(\Omega \left(\mathsf{E} - \mathsf{E}_\chi \right) \right) &= \log \left(\Omega (\mathsf{E}) \right) - \mathsf{E}_\chi \ \frac{\mathrm{dlog}(\Omega)}{\mathrm{dE}} + \dots \\ \mathrm{but} \ \left. \frac{\partial \log(\Omega)}{\partial \mathsf{E}} \right|_{\mathsf{V},\mathsf{N}} = \beta \end{split}$$



Canonical Partition Function

• Therefore: $P_{\chi} \propto \exp(-\beta E_{\chi})$

Canonical (or Boltzmann) distribution law



- Proportionality constant determined by normalisation condition: $\sum_{\chi} P_{\chi} = 1$
- Hence: $P_{\chi} = \frac{1}{Q} \exp(-\beta E_{\chi})$ with $Q = \sum_{\chi} \exp(-\beta E_{\chi})$
 - Q is referred to as the "canonical partition function"
 - It is a function of N, V, T (or equivalently N, V, β)
 - If the partition function of a system is known, one can derive any thermodynamic quantity!

Thermal Energy in the Canonical Ensemble

Consider for example the internal/thermal energy:

$$U = \langle E \rangle = \sum_{\chi} E_{\chi} P_{\chi}$$

$$\langle E \rangle = \frac{\sum_{\chi} E_{\chi} exp(-\beta E_{\chi})}{\sum_{\zeta} exp(-\beta E_{\zeta})}$$
 Probability of microstate χ \Rightarrow Canonical partition function, Q

$$\langle \mathsf{E} \rangle = \frac{1}{\mathsf{Q}} \sum_{\chi} \frac{\partial \exp \left(-\beta \, \mathsf{E}_{\chi} \right)}{\partial \left(-\beta \right)} \ = -\frac{1}{\mathsf{Q}} \frac{\partial}{\partial \beta} \sum_{\chi} \exp \left(-\beta \, \mathsf{E}_{\chi} \right) \ \Rightarrow$$

$$\langle \mathsf{E} \rangle = -\frac{\partial \log \mathsf{Q}}{\partial \beta} \bigg|_{\mathsf{N},\mathsf{V}}$$

Calculation of Entropy

 Gibbs provided a general equation for entropy, valid in fact for any ensemble:

$$S = -k_{B} \sum_{\chi} P_{\chi} \log(P_{\chi})$$

 In the microcanonical ensemble, this simplifies to Boltzmann's formula! (slide 19)



Josiah Willard Gibbs

Mark: prove statement above...

 $P_{\chi} = \frac{1}{\Omega} \exp(-\beta E_{\chi})$ In the canonical ensemble:

$$S = -k_{B} \frac{1}{Q} \sum_{x} exp(-\beta E_{x}) [-\beta E_{x} - log(Q)]$$

Entropy and Helmholtz Free Energy

Continuing with entropy in the canonical ensemble...

$$S = -k_{B} \frac{1}{Q} \sum_{x} exp(-\beta E_{x}) [-\beta E_{x} - log(Q)] \implies$$

$$S = -k_{B} \frac{1}{Q} \left[-\beta \sum_{\chi} E_{\chi} \exp(-\beta E_{\chi}) - Q \log(Q) \right] \Rightarrow$$

$$S = -k_{B} \left[-\frac{1}{k_{B}T} U - \log(Q) \right] = \frac{U}{T} + k_{B} \log(Q) \implies$$

$$TS = U + k_B T \log(Q)$$

• But from thermodynamics we know: A = U - T S (slide 6)

... therefore:
$$A = -k_B T \log(Q)$$

Pressure in the Canonical Ensemble

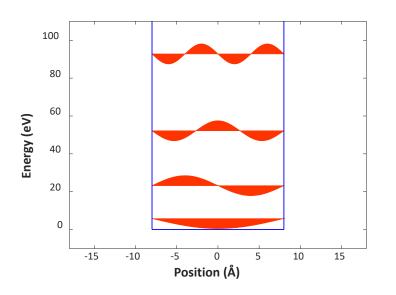
Pressure is now easy to calculate, since:

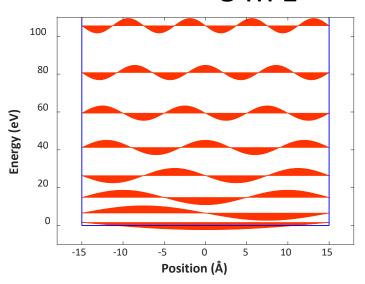
we easy to calculate, since:
$$\frac{\partial X}{\partial V}\Big|_{N,T} = -p$$

$$\Rightarrow p = k_B T \frac{\partial \log(Q)}{\partial V}$$

• The volume doesn't appear explicitly in Q, but it's there... It affects the energy levels!

– e.g. for a quantum particle in a 1D box:

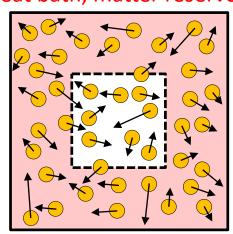




Grand Canonical Ensemble (µ, V, T)

- This ensemble comprises microstates with fixed volume
- Energy and number of molecules can fluctuate and the system is at thermodynamic equilibrium with
 - a heat bath kept at temperature T
 - a matter reservoir kept at chemical potential μ
- Heat bath & matter reservoir >> system

heat bath, matter reservoir



Matter exchange Heat exchange

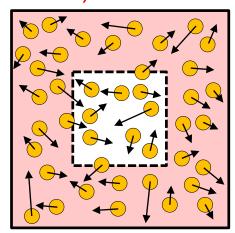
One can apply a similar reasoning as that for the canonical ensemble to prove that...

 The probability in the Grand Canonical Ensemble is:

$$P_{\chi} = \frac{1}{\Xi} \exp(-\beta E_{\chi} + \beta \mu N_{\chi})$$

with
$$\Xi = \sum_{\chi} \exp(-\beta E_{\chi} + \beta \mu N_{\chi})$$

heat bath, matter reservoir



Matter exchange Heat exchange

- $-\Xi$ is referred to as the "grand canonical partition function"
- It is a function of μ , V, T (or equivalently μ , V, β)

Number of Particles in Grand Canonical Ensemble

- In the microcanonical and canonical ensembles, the number of particles (molecules) was fixed.
- Not the case in grand canonical! How can we find the average number of particles?

$$\langle N \rangle = \frac{1}{\Xi} \sum_{x} N_{x} \exp(-\beta E_{x} + \beta \mu N_{x}) \implies$$

$$\langle N \rangle = \frac{1}{\Xi} \sum_{\chi} \frac{\partial \exp(-\beta E_{\chi} + \beta \mu N_{\chi})}{\partial (\beta \mu)} \Rightarrow$$

$$\langle N \rangle = \frac{1}{\Xi} \frac{\partial}{\partial (\beta \mu)} \sum_{\chi} \exp(-\beta E_{\chi} + \beta \mu N_{\chi}) = \frac{\partial \log(\Xi)}{\partial (\beta \mu)}$$

Entropy and Pressure in Grand Canonical Ensemble

Let us calculate the entropy from Gibbs' formula:

$$\begin{split} S &= -k_{_B} \sum_{\chi} P_{\chi} log(P_{\chi}) \implies \\ S &= -k_{_B} \sum_{\chi} P_{\chi} \Big[-log(\Xi) - \beta \, E_{\chi} + \beta \mu N_{\chi} \Big] \implies \\ S &= -k_{_B} \Big[-log(\Xi) - \beta \, \langle E \rangle + \beta \mu \langle N \rangle \Big] \implies \\ Recall: \\ \beta &= (k_{_B}T)^{-1} & TS &= k_{_B} T log(\Xi) + U - \mu N \end{split}$$

$$pV = k_B T log(\Xi)$$

Take Home Messages

- Statistical Mechanics is all about deriving expressions for familiar macroscopic thermodynamic variables from microscopic laws (e.g. Schrödinger's equation or Newton's law of motion)
- Two fundamental postulates:
 - Principle of equal weights...
 - Ergodic hypothesis...
- Averaging makes use of ensembles. Commonly used:
 - Microcanonical (N, V, E)
 - Canonical (N, V, T)
 - Grand canonical (μ , V, T)
- Macroscopically, all ensembles give "same answers"...